A Machine Learning Approach for Automated Sunquake Detection

Vanessa Mercea

Technical University of Cluj-Napoca

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Work with: A.R. Paraschiv, D.A. Lăcătuș, A. Mărginean, D. Beșliu-Ionescu

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Definition

Sunquakes are progressive circular waves observed on the fotosphere, produced by solar flares.

Interest

Manual detection is often laborious. Manifestation circumstances are not entirely known. Several detection methods are available but none are automatic.

Sunquakes

Helioseismic Holography



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Time Distance





Movies/Wave Detection



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Data Morphology



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Image: A matrix

Data Morphology



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imbalance: on average, 4% of all individual data samples are SQ

complexity: intensity and variety of SQ signature patterns

artefacts: "eye", AR shadow signatures, intense solar storms

Artefacts





- sources: SC23 (1), SC24 (2)
- MDI (3) and HMI (4) Dopplergram download (batch script)
- coordinate conversion

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- Holography method application (5)
- two obtained datsets (SunquakeNet;DOI10.34740):
 - acousic emission maps in FITS format: 53 (15 + 38)
 - grayscale 2D images in JPEG format:
 - ★ positive class: 845 (SC23: 205 + SC24: 640)
 - * negative class: 13.055 (SC23: 3891 + SC24: 9164)

Reconstruction-based learning

unsupervised AutoEncoder feature extraction feature classification

Contrastive learning

unsupervised and/or supervised CL feature extraction feature classification

Object detection

region proposal candidate region classification

Reconstruction-based learning



Figure: AutoEncoder architecture

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Reconstruction-based experiments (AE & VAE)

- AutoEncoder
 - fully black reconstructions (ls \leq 512)
 - poor classification results
- Variational AutoEncoder (VAE) (6)
 - attempt to capture distinct characteristics
 - almost fully black reconstructions (ls \leq 512)
 - little improvement in classification



Figure: AE vs. VAE architecture

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Reconstruction-based experiments (Log-Cosh VAE)

- Log-Cosh VAE (7)
 - ▶ \uparrow values: L_1 loss + \downarrow values: L_2 loss ⇒ Log-cosh loss
 - attempt to mitigate the impact of noise
 - poor reconstructions (ls \leq 512)
 - consistent improvement in classification



Figure: Log-cosh loss plots

Reconstruction-based learning



Figure: Log-Cosh VAE reconstructions

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Contrastive Learning



Figure: self-supervised CL (left), supervised CL (right) applied to SunquakeNet

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- self-supervised CL (8)
 - results similar to Log-Cosh VAE
 - sunquake signatures are not consistently captured
- supervised CL (9)
 - attempt to capture distinct characteristics
 - tedious and unstable training due to imbalance
- self-supervised CL using upsampling \Rightarrow supervised CL with a weighted loss (10):
 - attempt to mitigate the impact of imbalance
 - significantly improved results

Object Detection

Steps

identify multiple candidate object regions classify whether or not each region is a Sunquake return regions classified as Sunquake



Augmentation Methods

- Custom domain-specific Augmentations:
 - Altered Random Erase:
 - * forces the network to be attentive to all areas inside the image
 - \star replaces the typically used Random Crop
 - * decreases the probability of occluding a Sunquake
 - Solarized Low Pass Filter:
 - \star enhances high frequency signals and fades the others out
 - * amplifies details for some Sunquake signatures
 - ► Time Based Mixing:
 - \star combines successive grayscale frames into a single 3D image
 - * maintains the sequence property of data
- General Augmentations:
 - Geometric transforms:
 - ★ flips (horizontal and vertical)
 - * rotations (90, 180, 270)

Altered Random Erase



Figure: Visualization of the effect of the Altered Random Erase Augmentation

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Solarized Low Pass Filter



Figure: Visualization of the effect of Solarized Low Pass Filter Augmentation

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Figure: Visualization of the effect of Time Based Mixing Augmentation



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Loss Weighting

Regular supervised CL loss

$$l_{i} = \sum_{i \in I} \frac{-1}{|P(i)|} \sum_{p \in P(i)} \log \frac{e^{\frac{sim(z_{i}, z_{p})}{\tau}}}{\sum_{k=1}^{2N} mask_{[k \neq i]} e^{\frac{sim(z_{i}, z_{k})}{\tau}}}$$
(1)

Weighted supervised CL loss

$$l_{i} = \sum_{i \in I} \frac{-w_{y_{i}}}{|P(i)|} \sum_{p \in P(i)} log \frac{e^{\frac{sim(z_{i}, z_{p})}{\tau}}}{\sum_{k=1}^{2N} mask_{[k \neq i]} e^{\frac{sim(z_{i}, z_{k})}{\tau}}}$$
(2)
$$w_{y_{i}} = \frac{1 - \beta}{1 - \beta^{N_{y_{i}}}}$$
(3)

- the number of positive image samples is increased for unsupervised methods
- five extra copies are generated for each image (flips and rotations)
- may impose a transformation bias to the model if used in supervised methods

SMOTE

Definition

synthetic minority oversampling technique



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Sunquake time prediction

- CL Model with shuffle bias
 - \blacktriangleright frame level shuffling \Rightarrow shared Active Regions between the training and validation sets
 - training specifications (two-step):
 - * supervised CL, 50 epochs: DenseNet-121, temperature 0.1, positive upsampling
 - \star embeddings classification: various classifiers trained using SMOTE
- CL Model with no bias
 - event level shuffling \Rightarrow unique Active Regions in the training and validation sets
 - training specifications (three-step):
 - * self-supervised CL, 500 epochs: ResNet-18, positive upsampling
 - * supervised CL 100 epochs: weighted loss, temperature 0.07
 - \star embeddings classification: various classifiers trained using SMOTE

Sunquake location prediction

- Object Detection Model
 - faster R-CNN, 50 epochs, trained only on positive SC23 image data and regions

Process Diagram



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Classifier	K-NN (Bagging)		SVC (Poly)		SVC (RBF)		Logistic Regression		SGD	
Augmentation	None	SMOTE	None	SMOTE	None	SMOTE	None	SMOTE	None	SMOTE
Precision	0.97	0.80	0.97	0.98	0.99	0.94	0.97	0.91	0.91	0.89
Recall	0.54	0.96	0.96	0.97	0.86	0.99	0.98	0.99	0.94	0.93
F1-Score	0.55	0.86	0.97	0.97	0.91	0.96	0.98	0.95	0.94	0.93
Accuracy	0.94	0.95	0.99	0.99	0.98	0.99	0.99	0.99	0.98	0.98
Metrics Avg	0.750	0.892	0.947	0.977	0.935	0.97	0.98	0.96	0.942	0.932

Classifier	K-NN (Bagging)		SVC (Poly)		SVC (RBF)		Logistic Regression		SGD	
Augmentation	None	SMOTE	None	SMOTE	None	SMOTE	None	SMOTE	None	SMOTE
Precision	0.63	0.65	0.66	0.84	0.49	0.59	0.64	0.64	0.54	0.62
Recall	0.54	0.54	0.54	0.54	0.50	0.54	0.54	0.54	0.54	0.54
F1-Score	0.55	0.55	0.55	0.55	0.49	0.55	0.55	0.55	0.54	0.55
Accuracy	0.93	0.93	0.93	0.94	0.93	0.92	0.93	0.93	0.89	0.93
Metric Avg	0.662	0.667	0.67	0.715	0.605	0.65	0.665	0.665	0.627	0.66

Table: Macro Average performance of different classifiers over embeddings produced by the CL model with **top: shuffle bias**, **bottom: no bias**, trained with and without SMOTE augmentation, for the test data in **SC23 & SC24** (2622 negative and 186 positive samples)

Prediction Distribution



Figure: Predictions of the *left*: shuffle bias model; *right*: unbiased model, for the test data in SC23 & SC24, clustered by UMAP components

Event	Counts				Event	Counts			
Date	TP	FP	FN	GT	Date	ΤP	FP	FN	GT
1996-07-09 09:01	0	0	19	19	2012-07-04 09:47	2	0	17	19
2001-04-06 19:13	0	0	16	16	2012-07-06 13:26	2	0	15	17
2001-09-24 09:35	0	0	11	11	2013-11-08 04:20	2	0	18	20
2002-07-23 00:27	0	6	14	14	2015-03-11 16:11	2	0	14	16
2012-03-05 19:27	2	0	19	21	2015-09-28 14:53	2	0	18	20
2012-03-06 07:52	2	0	11	13					

Table: SVC (poly) predictions for embeddings produced by the unbiased CL model for the test data in SC23 & SC24 (2622 negative și 186 positive samples).

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Cosine distance characteristic analysis (unbiased model)



Figure: Cosine Distances computed between consecutive frames embeddings for the event at 2012-07-06 13:26 in the test set

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Mean characteristic analysis (unbiased model)



Figure: Means computed for frame embeddings for the event at 2012-07-06 13:26 in the test set

Cosine distance characteristic analysis (biased model)



Figure: Cosine Distances computed between consecutive frames embeddings for the event at 2012-07-06 13:26 in the test set

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problematic data cubes

AIA/RHESSI prediction analysis on additional datasets

Predictions for the problematic 2012.05.08 13:02 event



Figure: Position of candidate SQ signatures, at frames [180, 188), predicted positive for the 2012.05.08 13:02 event in the unlabeled dataset.

Prediction validation for the problematic 2012.05.08 13:02 event



Figure: Flaring Activity observed in the AIA 304 Å channel and the RHESSI location of seismic signatures for the 2012.05.08 13:02 event in the unlabeled dataset.

Predictions for the problematic 2011.12.30 03:03 event



Figure: Position of candidate SQ signatures, at frames [15, 22), predicted positive for the 2011.12.30 03:03 event in the unlabeled dataset.

Prediction validation for the problematic 2011.12.30 03:03 event



Figure: Flaring Activity observed in the AIA 94 Å channel and the RHESSI location of seismic signatures for the 2011.12.30 03:03 event in the unlabeled dataset.



- constructed and published two datasets of SC23 & SC24
 - strongly imbalanced
 - with artefacts
 - complex
- implemented multiple ML detection methods:
 - AutoEncoders (AE, VAE, Log-Cosh VAE)
 - CL (self-supervised, supervised, with weighted loss)
 - Object Detection (Faster R-CNN)
 - Various classifiers
- implemented 3 custom domain-specific augmentations:
 - Altered Random Erasing
 - Solarized Low Pass Filter
 - Time Based Mixing

AE results are poor, but a level of clustering was observed at event level

distinct CL models provide consistent results

distinct CL models capture a low similarity between Sunquake transition frames and their neighbors

correlations are found between predicted lower acoustic emission sources and Solar eruptions accompanied by high-energy X-Ray emissions Time and location prediction components are separated improve CL models with explainability techniques

Models predict Sunquakes of a too short duration mix more than 3 frames in the Time Based Mixing augmentation

Spurious correlations are not addressed

create a more fine-grained separation of signature pattern classes

Thank you!

Don't hesitate to reach out if you have any questions or ideas related to this topic!

Mercea.Fl.Vanessa@student.utcluj.ro / mercea.vanessa@yahoo.com

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