



Applications of Type I diffracted photons
in astronomical imaging

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Summary

- **Introduction**
- **Modified Thomas Young's double-slit experiment**
- **QCE image sensor**
- **Calibrations in optical microscopy**
- **Measurements with Type I diffracted photons in astronomical imaging**
- **Insensitivity of QCE image sensor to seeing conditions**
- **Conclusions**

Introduction (1)

- **Christiaan Huygens (1690)**
- **Isaac Newton (1704)**
- **Thomas Young (1803); August Fresnel (1815)**
- **Albert Einstein (1905); Arthur H. Compton (1923)**

Introduction (2)

Optical lithography

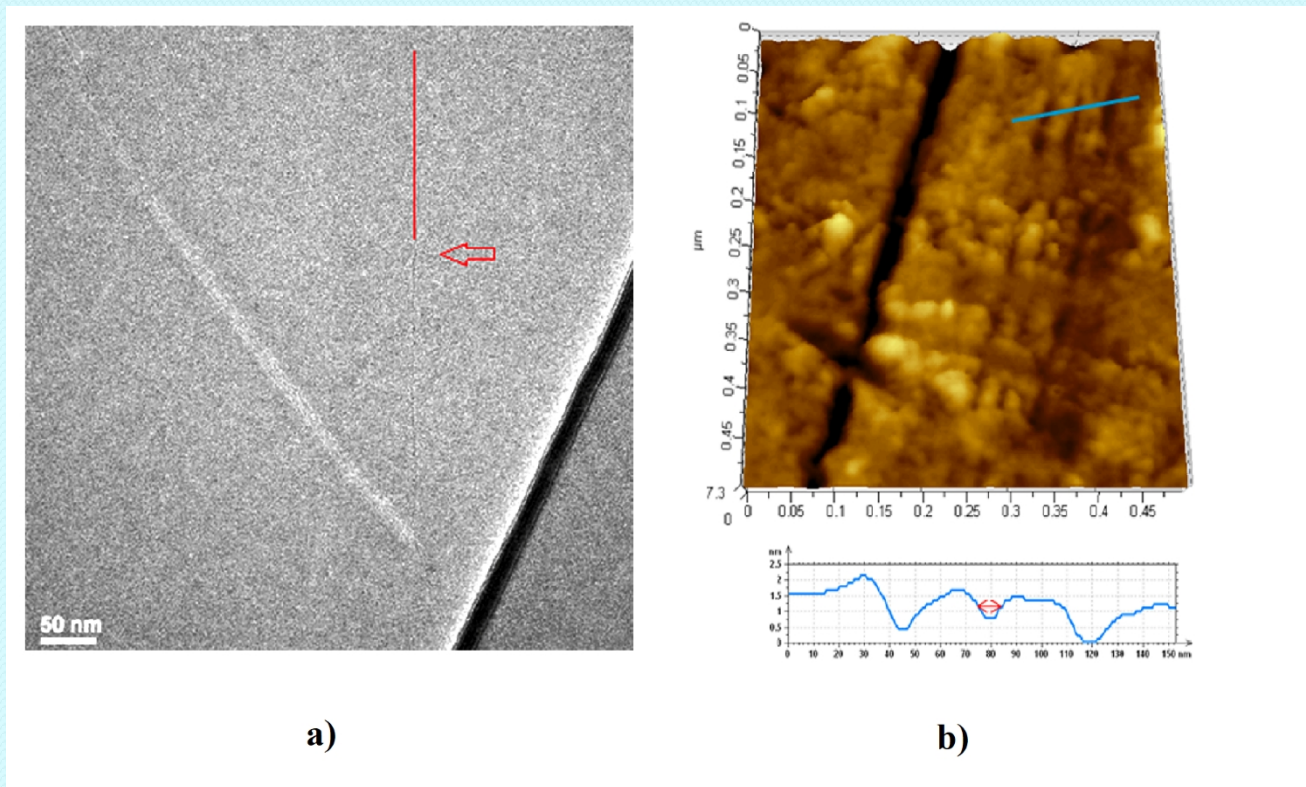
The resolution of the smallest printable feature is calculated by Rayleigh scaling equation:

$$D = k_1 \lambda / NA$$

where k_1 is a process parameter, λ is the light wavelength and NA is the numerical aperture of the optical system. With the parameters $k_1 = 0.35$, $\lambda = 650 \text{ nm}$, $NA = 0.6$, we obtain **D = 379 nm**.

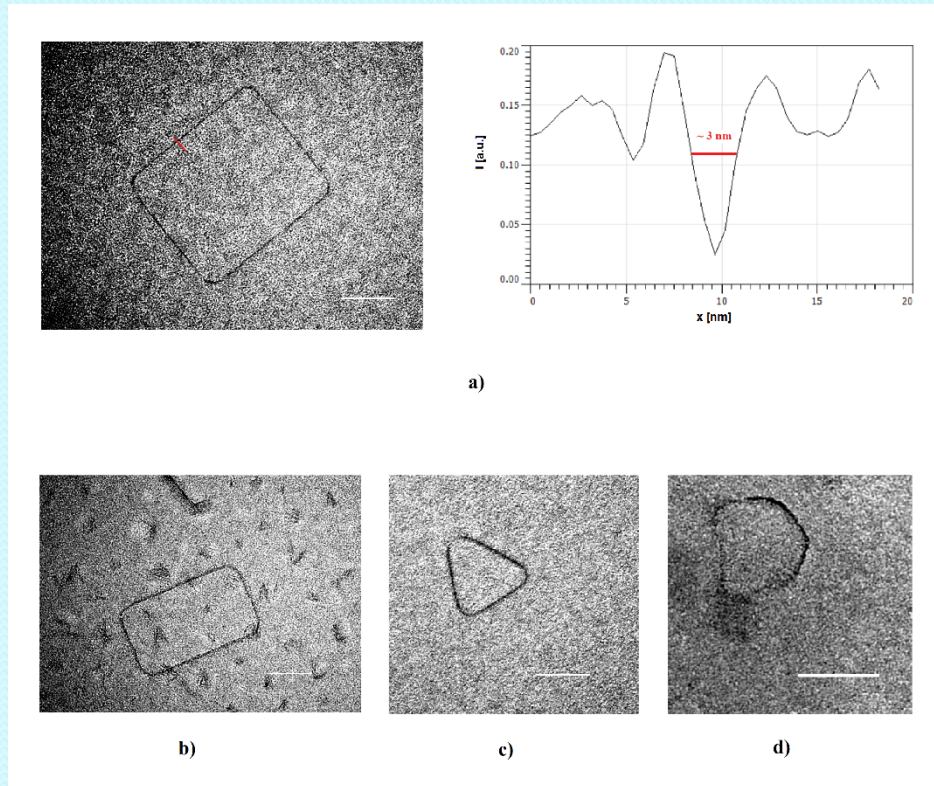
The quantum confinement effect (QCE) applied in Quantum Optical Lithography broke the diffraction limit.

Introduction (3)- Quantum Optical Lithography



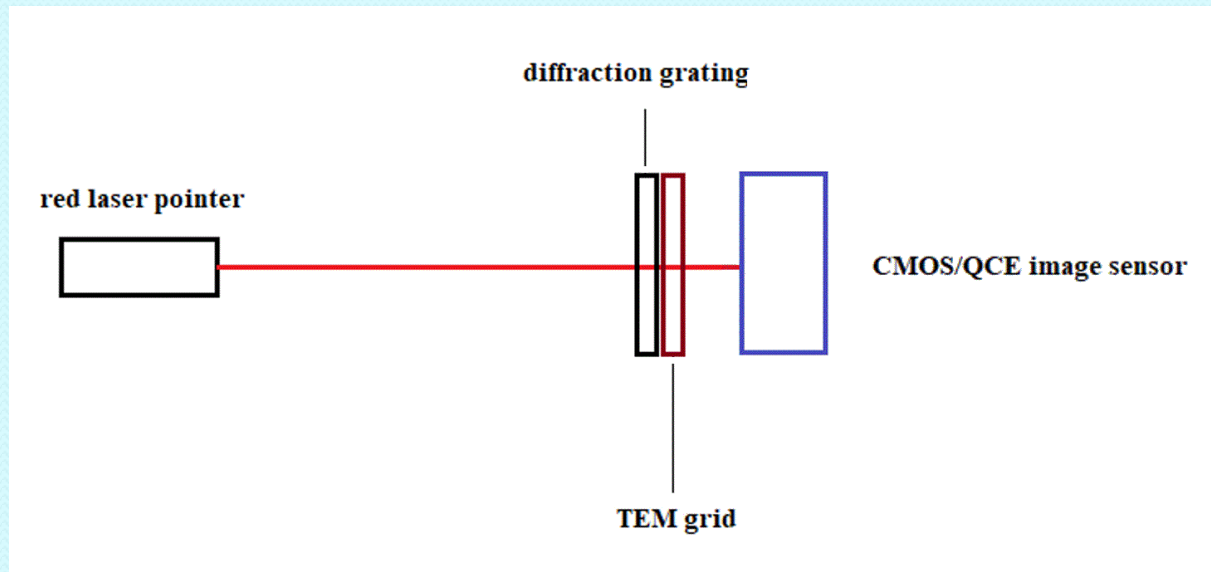
(a) TEM image of single 1 nm line. Ref. *Optics & Laser Technology* **60** (2014) 80–84; (b) AFM topography image of 2 nm parallel lines on sample surface. The cross-sections of 2 nm lines reveal line depths up to 1.5 nm – as deep as the tip can reach - and a width of 4-8 nm, broadened by the AFM tip. The thick line situated on the left side in both images represents a 20 nm marker line. Ref. *Optics Communications* **291** (2013) 259–263.

Introduction (4)- Quantum Optical Lithography



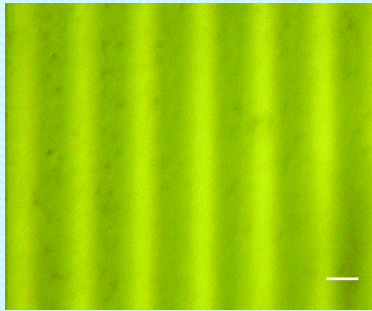
TEM images of different geometrical shapes written on Si_3N_4 TEM grids covered by resist: (a, b) rectangles written at 560 nW (scale bar, 100 nm); (c) triangle written at 510 nW. (scale bar, 100 nm); (d) "D" letter written at 510 nW (scale bar, 50 nm). Ref. *J. Micro/Nanolith. MEMS MOEMS*, **18**(2) (2019) 02050.

Modified Thomas Young's double-slit experiment (1)

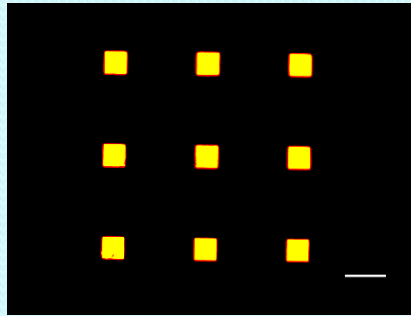


Experimental system used in multiple slit measurements.

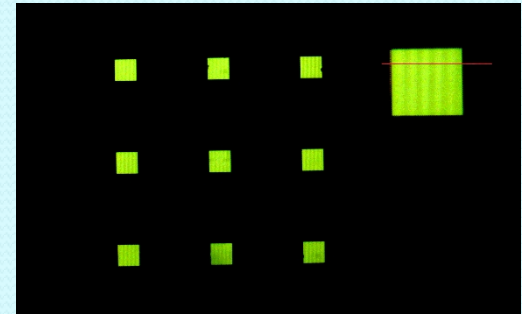
Modified Thomas Young's double-slit experiment (2)



a)



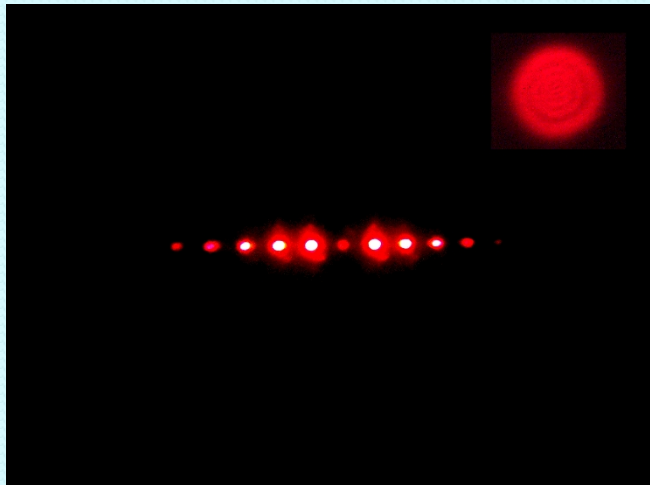
b)



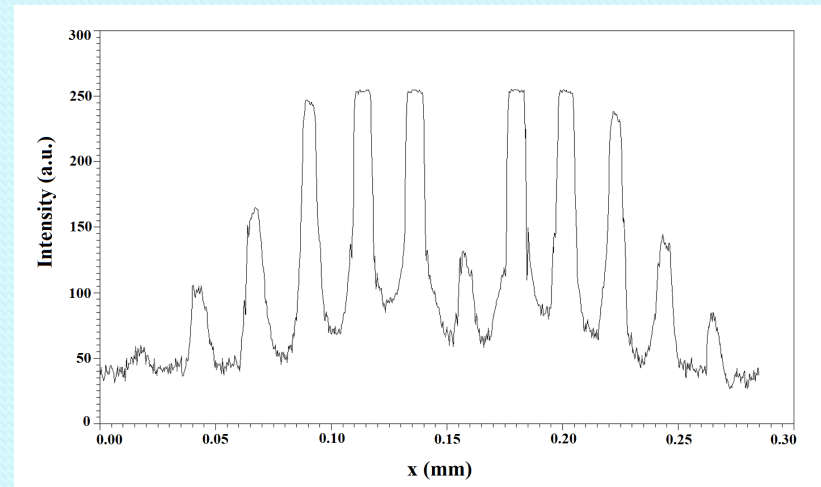
c)

Optical microscope images of: a) 50L grating, scale bar: 10 μm ; b) TEM grid, scale bar: 200 μm ; c) ensemble of 50L grating and TEM grid; insert: detailed image of a window.

Modified Thomas Young's double-slit experiment (3)



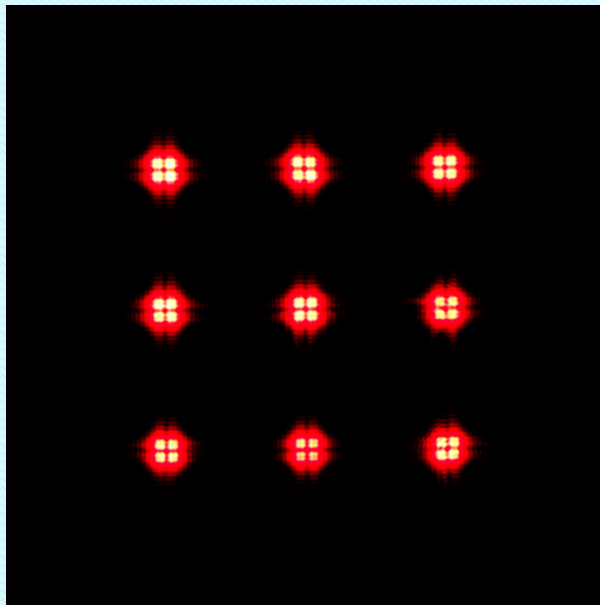
a)



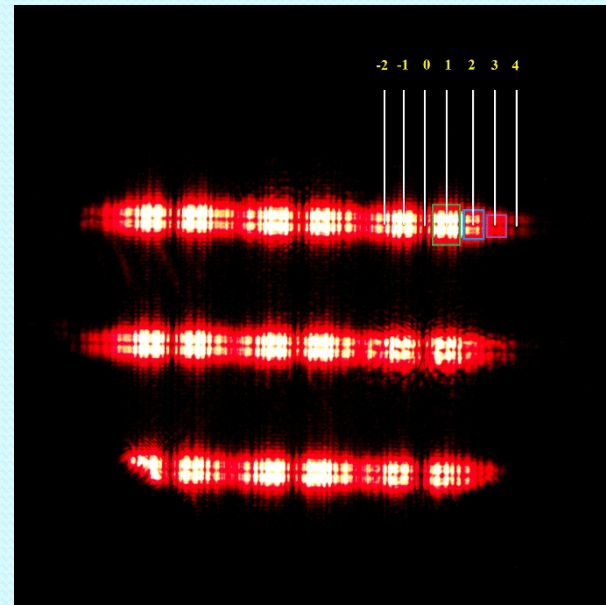
b)

a) Diffraction pattern of the laser pointer recorded in transmittance mode from 50L grating. The distance from the grating to the screen was $D_1 = 250$ mm. Insert: CMOS sensor image of TEM (4,0) cylindrical transversal mode of the laser beam; b) diffraction efficiency of grating (0, 1, 2, 3, 4, 5 orders).

Modified Thomas Young's double-slit experiment (4)



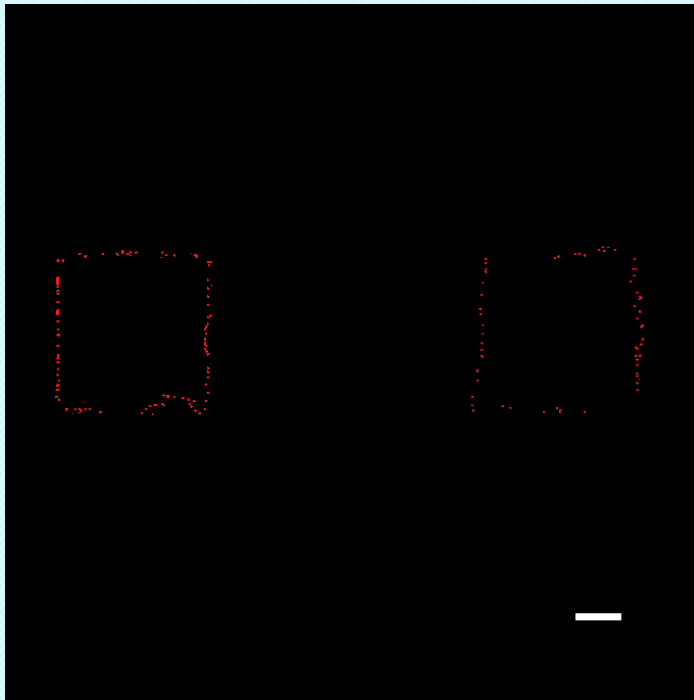
a)



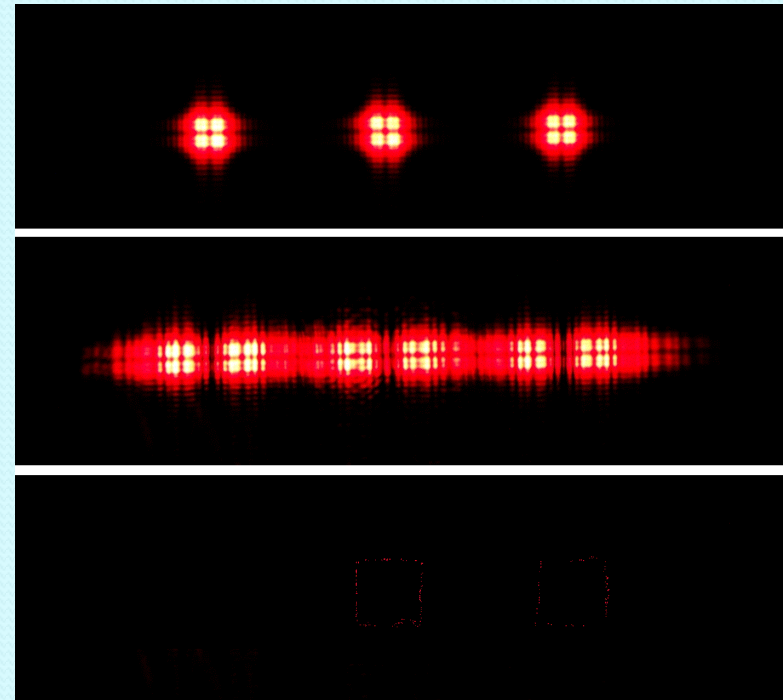
b)

a) CMOS image of the rectangular transverse mode TEM (1,1) recorded in transmittance mode from TEM grid; b) CMOS image of the diffraction patterns recorded in transmittance mode from 50L grating and TEM grid ensemble.

Modified Thomas Young's double-slit experiment (5)



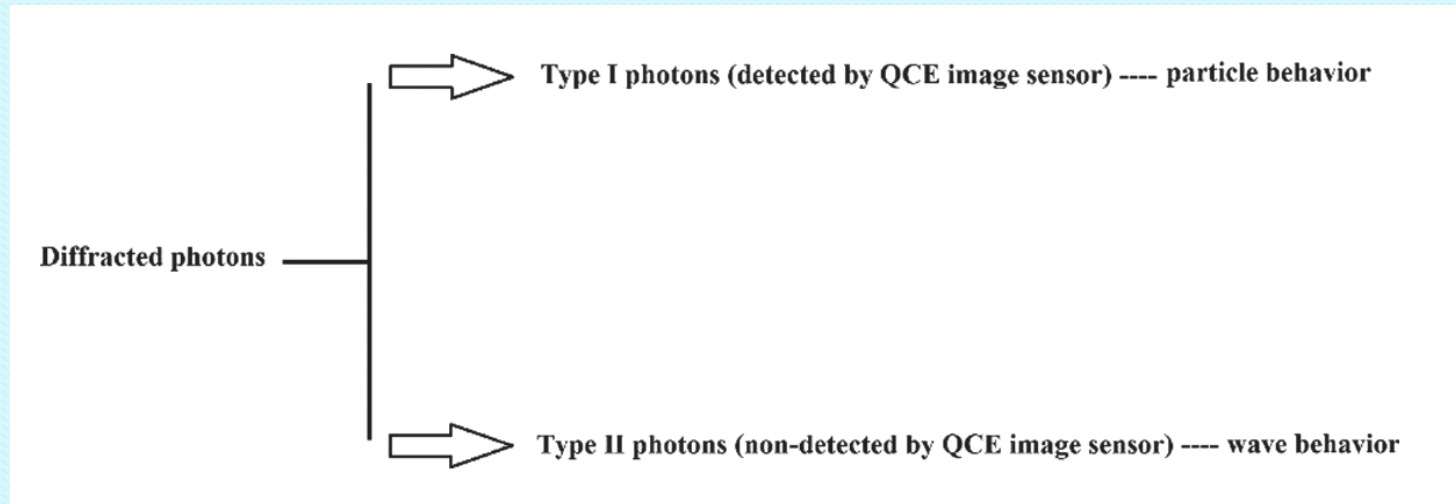
a)



b)

a) QCE image of the pattern recorded in transmittance mode from 50L grating and TEM grid sandwich; b) rows of CMOS images of 3 windows of TEM grid without (upper row) and with diffraction grating (middle row) and QCE image with diffraction grating (lower row).

Modified Thomas Young's double-slit experiment (6)



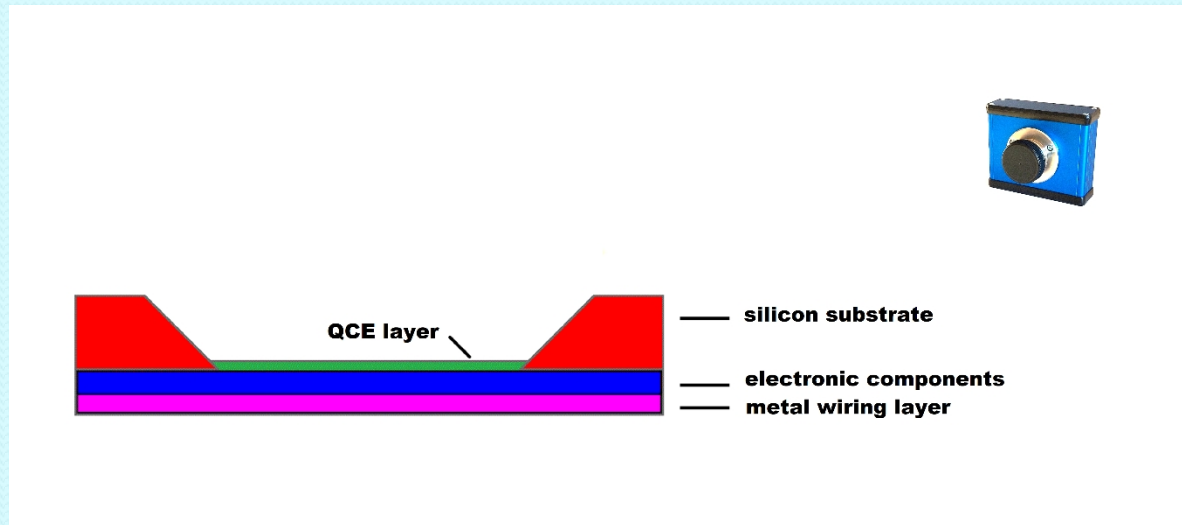
Classification of the diffracted photons.

Modified Thomas Young's double-slit experiment (7)

Table 1. Characteristics of the Diffracted Photon Types

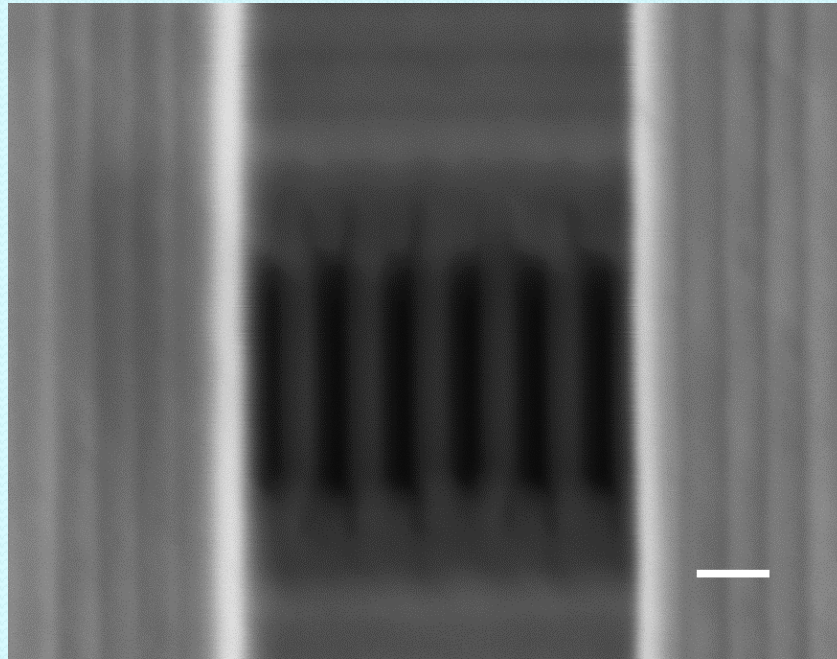
	Detector(s)	Linear Momentum of Diffracted Photons	Fluence (Energy Delivered per Unit Area)
Type I diffracted photons	QCE image sensor	Initial	$k \times F_0^{1/3}$ [16]
Type II diffracted photons	CMOS, CCD, photographic plate, eye, etc.	Modified	$F_0 - k \times F_0^{1/3}$

QCE image sensor (1)



Schematic cross section of BSI QCE image sensor.

QCE image sensor (2)



SEM image of 1Gb Winbond 25N01GVSFIT NAND memory.
Word lines have a width of 45 nm (scale bar: 100 nm).

QCE image sensor (3)

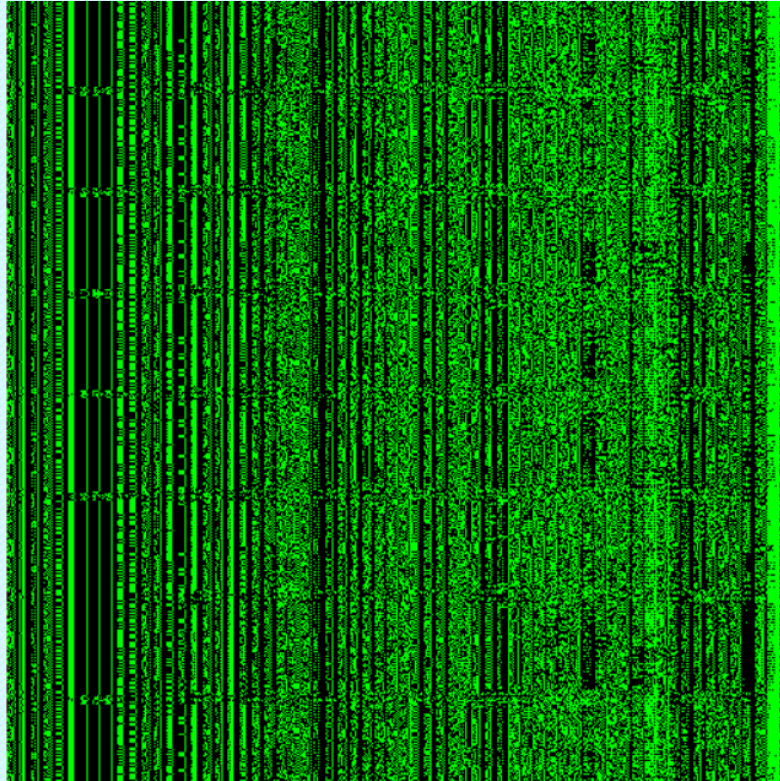


Image of a memory zone with 512 x 512 cells.

Calibrations in optical microscopy (1)

Objective	40x
NA	0.65
Diffraction limit*(nm)	619
Image resolution with A5314UPB 1.4MPixel CCD camera (nm/pixel)	168
Image resolution with QCE ST01D9 1GPixel image sensor (nm/pixel)	8.3
Magnification of QCE ST01D9 1GPixel image sensor vs diffraction limit (x)	74

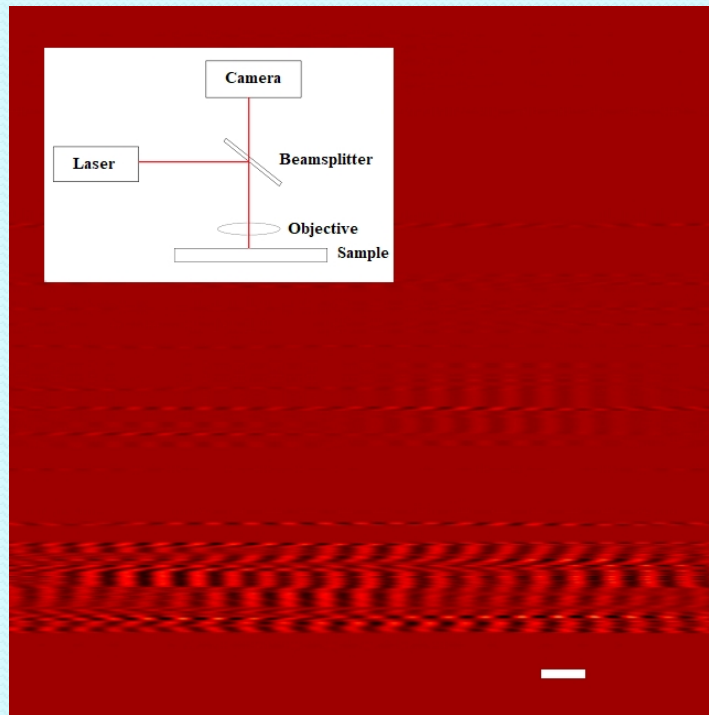
Table1. Characteristics of the system composed of optical microscope and camera.

* $R=0.61 \times \lambda/NA$, $\lambda=660\text{nm}$;

Pixel dimension: i) A5314UPB 1.4MPixel CCD camera: $4.65 \mu\text{m}$.

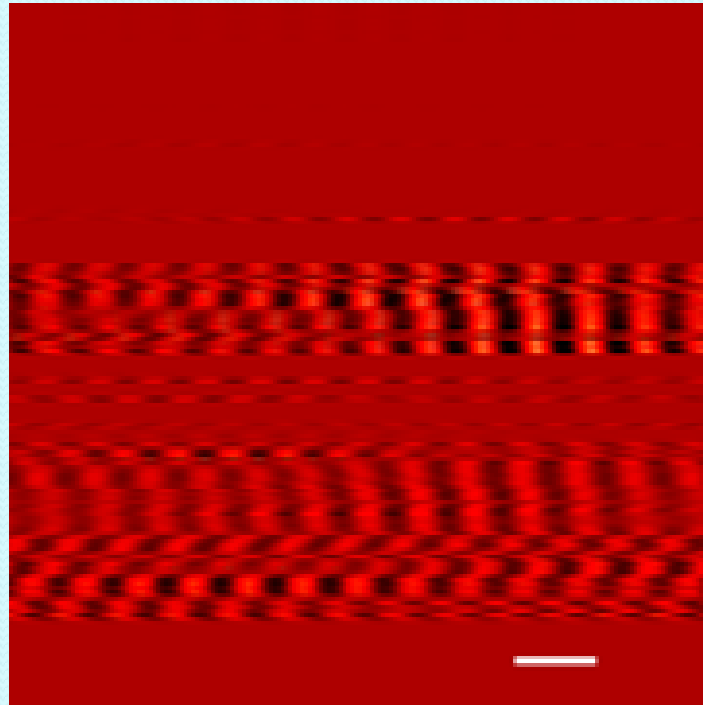
ii) QCE STM01D9 1GPixel image sensor: 60 nm , 76 times smaller than the pixel of A5314UPB 1.4MPixel CCD camera.

Calibrations in optical microscopy (2)



Sensor image of CD-R, objective 40x, scale bar 2 μm .

Calibrations in optical microscopy (3)



Sensor image of Blu-Ray sample, objective 40x, scale bar: 500 nm.

Measurements with Type I diffracted photons in astronomical imaging (1)

Table 1. Characteristics of the potential astronomical applications of ST01D9 1GPixel QCE image sensor

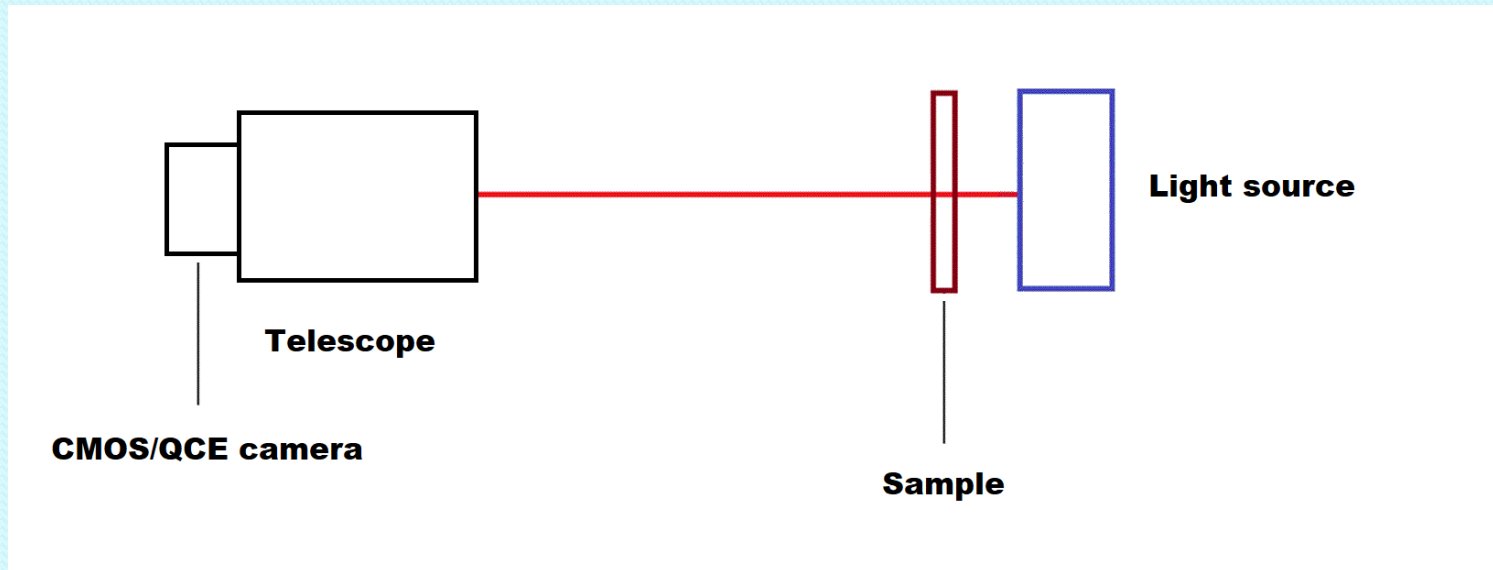
	Gran Telescopico Canarias	Hubble Space Telescope	J. Webb Space Telescope	N150/750 with 2X Barlow lens	Refractor 50/600
Φ (mm)	10400	2400	6500	150	50
Focal ratio (F/D)	16.33	24	20.2	10	12
Optimal pixel dimension* (μm)	9.96	14.6	12.3	6.1	7.32
Diffraction-limited resolution (arcsecond)**	0.012	0.052	0.019	0.838	2.516
Amplification factor*** (x)	120	175	148	73	88
Theoretical angular resolution with ST01D9 1GPixel QCE image sensor (arcsecond)	0.000100	0.000297	0.000128	0.011402	0.028590

* Optimal pixel dimension at $\lambda = 500 \text{ nm}$: $d(\mu\text{m}) = 0.61x(F/D)$.

** Rayleigh formula at $\lambda = 500 \text{ nm}$.

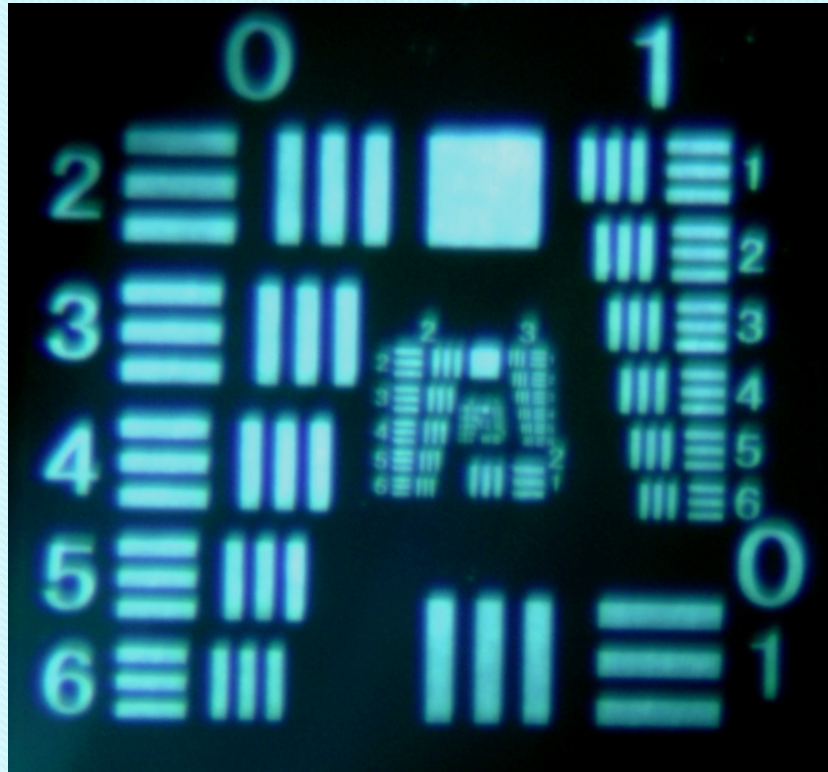
*** Optimal pixel dimension divided by mean pixel dimension (83 nm) of ST01D9 1GPixel QCE image sensor.

Measurements with Type I diffracted photons in astronomical imaging (2)



Experimental setup used in indoor measurements.

Measurements with Type I diffracted photons in astronomical imaging (3)



Width of a line, in millimetres,

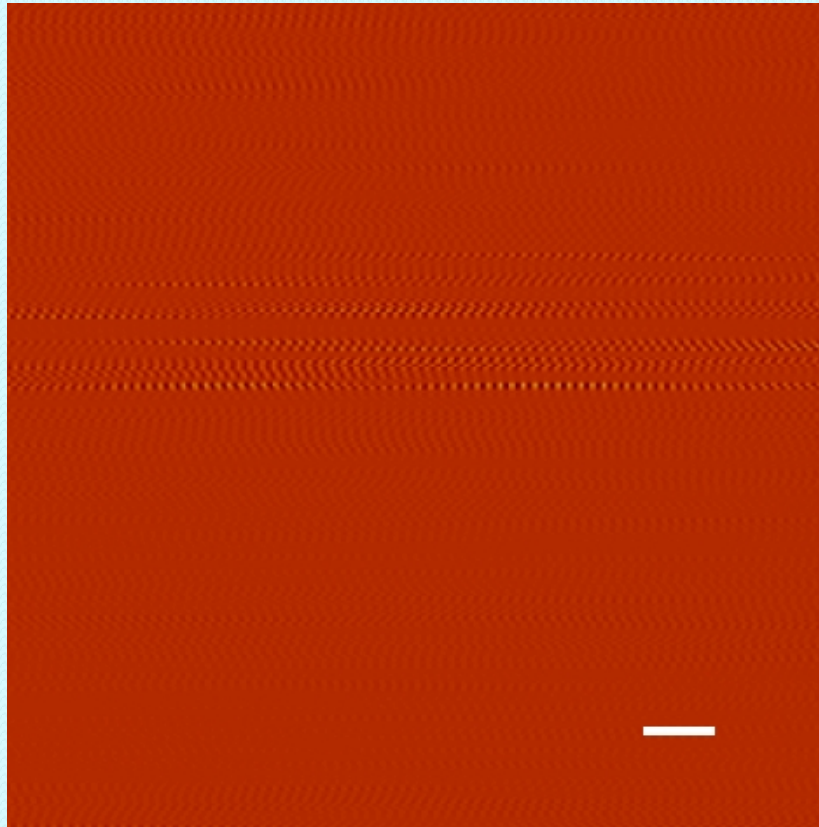
$$L(G; E) \text{ (mm)} = 2^{-G-(E+5)/6}$$

where G is the group number and E is the element number within the group.

The measured angular resolution of $L(3,5)$ is $\theta = 0.804$ arcsecond, close to Rayleigh diffraction limit $\theta = 0.838$ arcsecond and Dawes' limit $\theta = 0.773$ arcsecond.

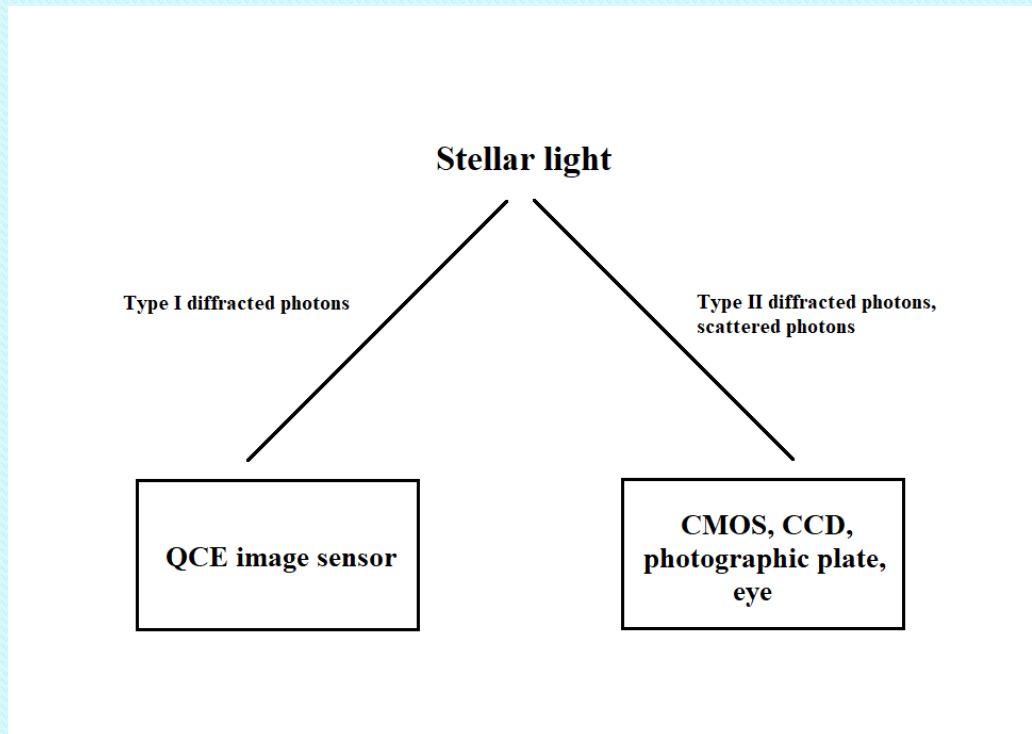
CMOS camera image of 1951 USAF resolution test chart situated at 10 m from a N150/750 telescope with 2X Barlow lens.

Measurements with Type I diffracted photons in astronomical imaging (4)



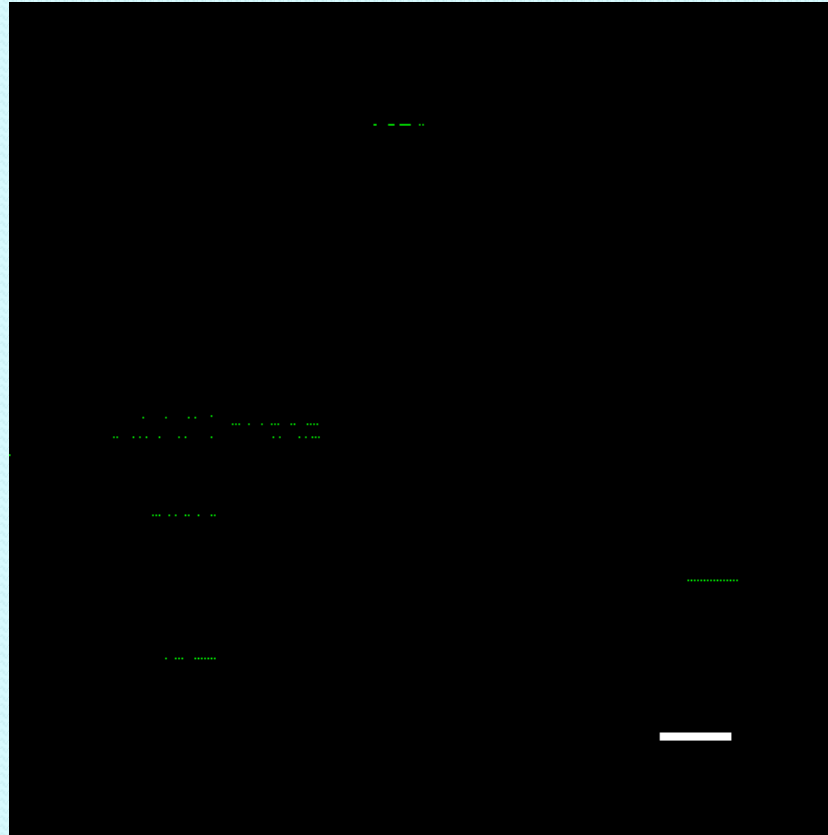
The picture captured by QCE camera of a CD-ROM sample situated at 10 m from a N150/750 telescope with 2X Barlow lens (scale bar: 200 mas).

Measurements with Type I diffracted photons in astronomical imaging (5)



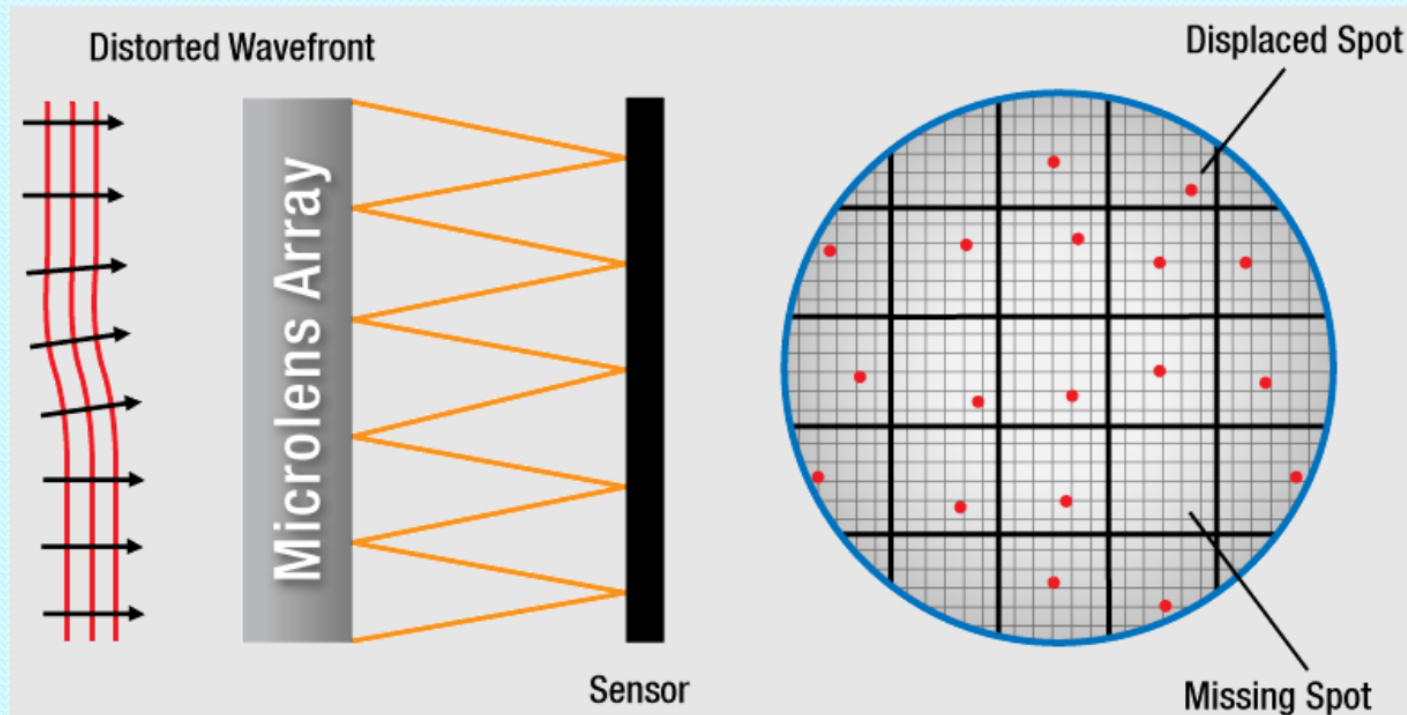
Classification of the light detectors versus photons types.

Measurements with Type I diffracted photons in astronomical imaging (6)



Multiple passages of Sirius (α Canis Majoris) captured by a QCE camera connected to N150/750 telescope with 2X Barlow lens (scale bar: 500 mas).

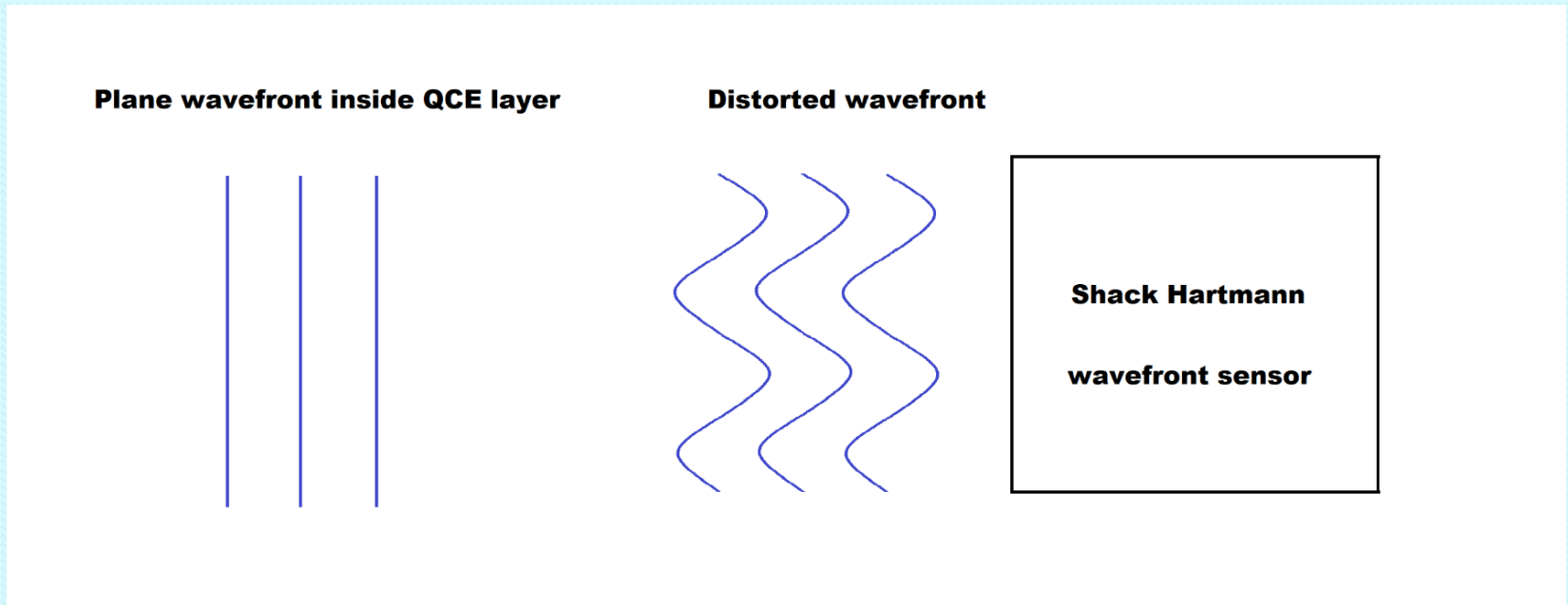
Insensitivity of QCE image sensor to seeing conditions (1)



Shack Hartmann wavefront sensor.

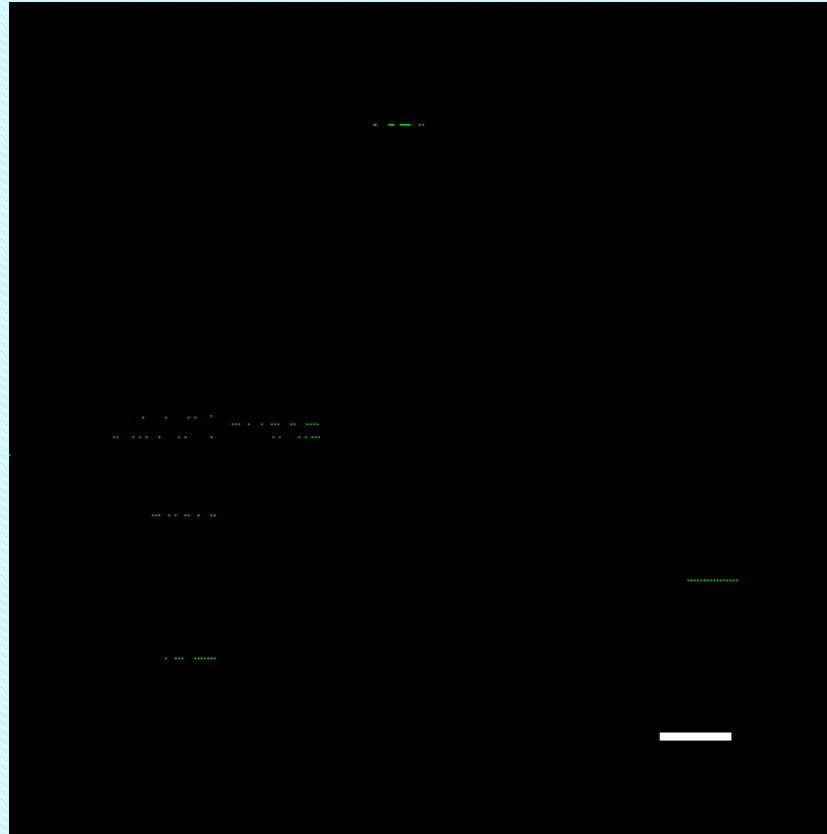
Ref. https://www.thorlabs.com/newgrouppage9.cfm?objectgroup_id=5287

Inensitivity of QCE image sensor to seeing conditions (2)



Wavefronts inside QCE layer and before Shack Hartmann wavefront sensor.

Insensitivity of QCE image sensor to seeing conditions (3) – Earth observations



Multiple passages of Sirius (α Canis Majoris) captured by a QCE camera connected to N150/750 telescope with 2X Barlow lens (scale bar: 500 mas).

Insensitivity of QCE image sensor to seeing conditions (4) – deep space observations



Pillars of Creation images of: a) NASA's Hubble Space Telescope (visible domain) and
b) NASA's James Webb Space Telescope (near-infrared domain).

Conclusions

- ❑ A modified Thomas Young's double-slit experiment was realized with Quantum Confinement Effect (QCE) image sensor. Two types of diffracted photons have been discovered: (i) Type I with particle behavior and (ii) Type II with wave behavior.
- ❑ The angular resolution of telescope due to Type I diffracted photons was investigated. Indoor measurements of the angular resolution have reached 16.5 milliarcseconds, an improvement of 50 x towards diffraction limit for a N150/750 telescope with 2X Barlow lens.
- ❑ Imaging through the atmosphere is improved by QCE image sensor due Type I diffracted photons with particle behavior. Sirius (α Canis Majoris) passage measurements have indicated seeing conditions estimated at 11 milliarcseconds. A novel concept is proposed to eliminate the seeing effects of atmospheric turbulence and to impulse the deep space investigations (nebulae, exoplanets, galaxy clusters, etc.).



Thank you for your attention!